



Conventional EO Satellites vs. CubeSats

-

FLD, AI flood detection onboard NanoSATs

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FDL: Josh Veitch-Michaelis, Gonzalo Mateo-García, Lewis Smith, Silviu Oprea,
Yarin Gal, Guy Schumann, Atılım Güneş Baydin, Dietmar Backes

Content:

Part 1: Conventional EO Satellites vs. CubeSats

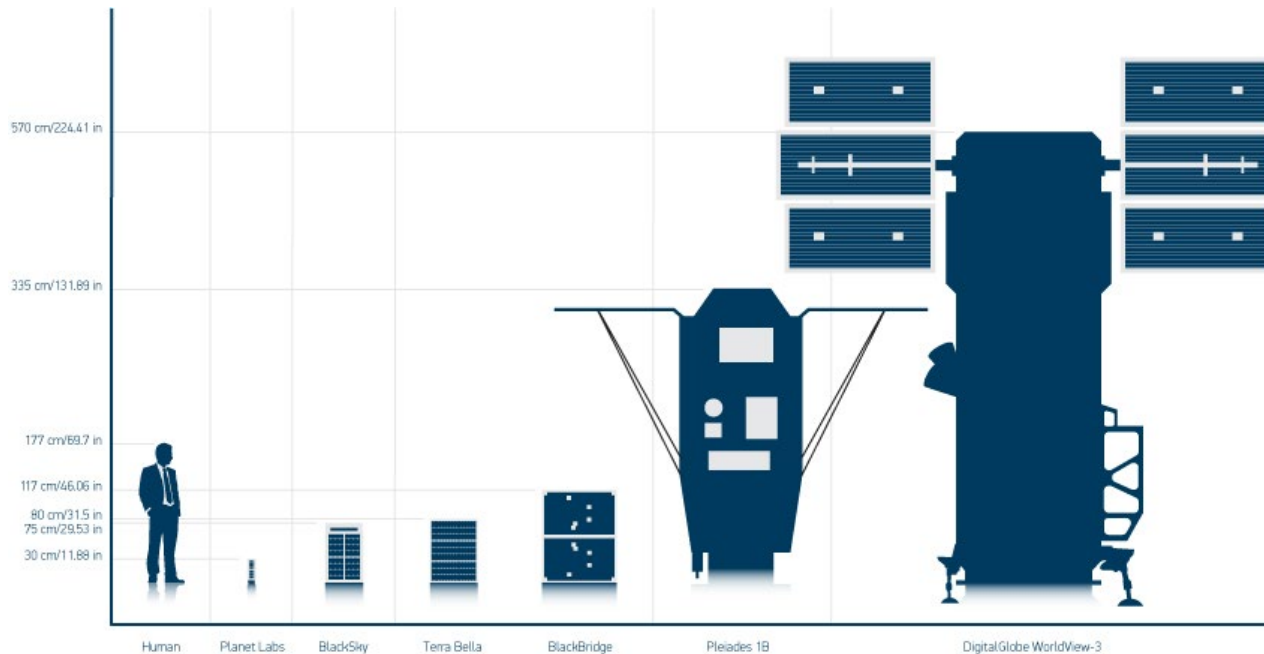
- a light hearted Overview
- Optical EO from LEO
- Basic Principles of Optical Spaceborne Imaging

Part 2: FDL-Europe - Disaster Prevention, Progress and Response (Floods)

- Flood Detection On Low Cost Orbital Hardware – Story and approach
- World Floods data set
- End to End machine learning approach
- Optimisation for deployment onboard the Satellite

Summary and Outlook

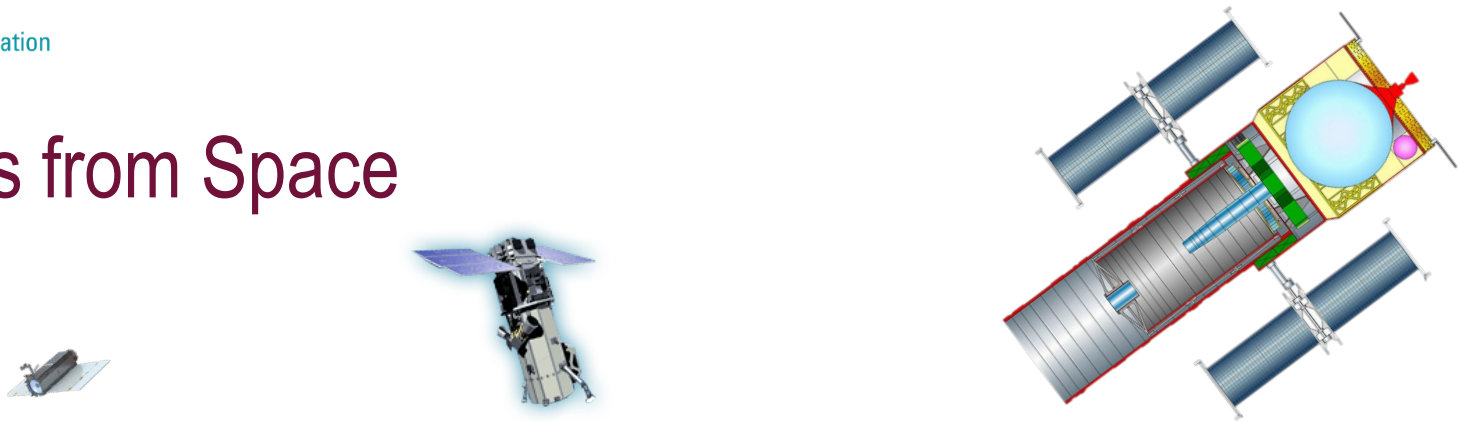
Conventional EO Satellites vs. CubeSats



Credit: Digital Globe 12/2016

...so where are the differences?

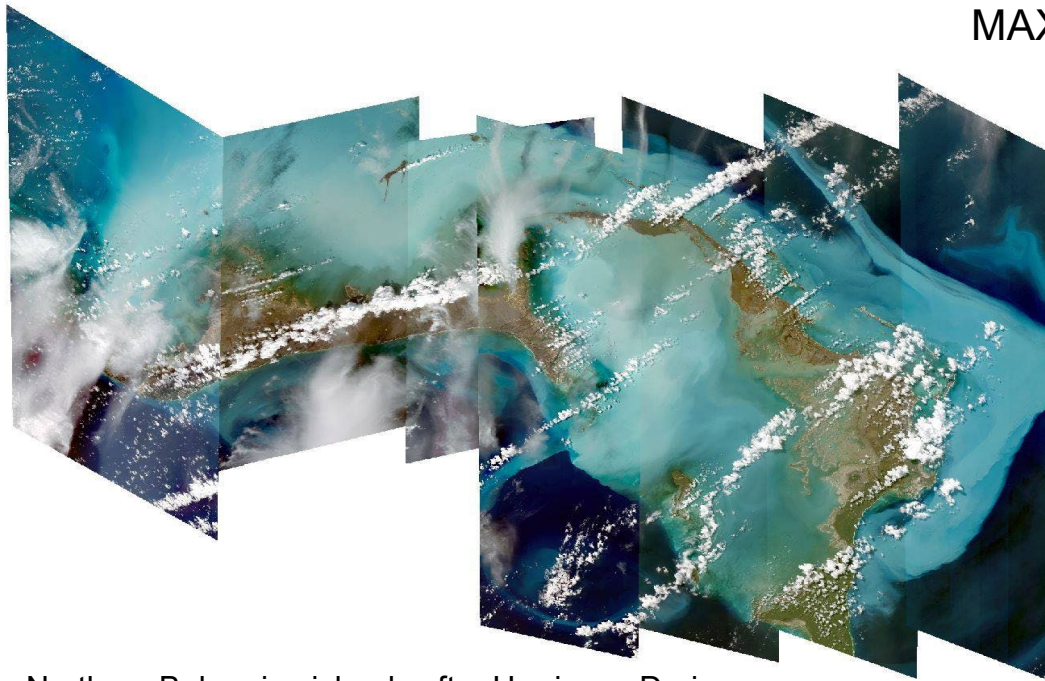
Pictures from Space



Spaceflight now 30/08/2019, Images of Semnan launch site Credit: Planet, Maxar and @realDonaldTrump

- Left: captured by Planet approx. 3m resolution (Dove, RapidEye or Skysat)
- Centre: Maxar/Digital Globe WorldView2 Satellite approx. 0.3m resolution
- Right: US intelligence image <https://twitter.com/realDonaldTrump/status/1167493371973255170> suspected KH11 type satellite (USA-244)

Very High-resolution Images by WorldView



MAXAR



WorldView 2 Satellite

- Multispectral images captured from a single orbit
- 29,900 km² in 7 separate scenes

Northern Bahamian islands after Hurricane Dorian

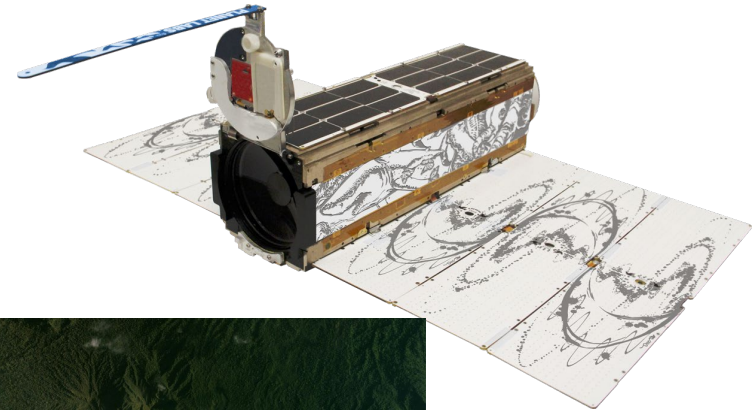
- Highly agile: body-pointing range of $\pm 40^\circ$ correspondent to 1355km FOR cross-track
- Pointing accuracy <500 m at image start and stop
- Large 2.2 TB on-board storage;
- Imagery is downlinked in X-band at 800 Mbit/s
- Theoretical 1.1 day revisit time

High-resolution Optical Images from CubeSats

Table 2
Preliminary assessment of the feasibility of Cubesat-based missions carrying different remote sensing technologies.

Technology	Feasibility assessment (feasible/problematic/ infeasible)	Justification
Atmospheric chemistry instruments	Problematic	Low sensitivity in SWIR-MIR because of limited cooling capability
Atmospheric temperature and humidity sounders	Feasible	e.g., GNSS radio occultation, hyperspectral millimeter-wave sounding
Cloud profile and rain radars	Infeasible	Dimensions, power
Earth radiation budget radiometers	Feasible	[63]
Gravity instruments	Feasible	[64]
High resolution optical imagers	Infeasible	Not enough resolution-swath, because limited space for optics and detectors
Imaging microwave radars	Infeasible	Limited power
Imaging multi-spectral radiometers (vis/IR)	Problematic	Limited imaging capability
Imaging multi-spectral radiometers (passive microwave)	Problematic	Limited imaging capability
Lidars	Infeasible	Limited power
Lightning imagers	Feasible	[30]
Magnetic field instruments	Feasible	[65]
Multiple direction/polarization radiometers	Problematic	Limited dimensions for receiver electronics
Ocean color instruments	Feasible	[4]
Precision orbit	Feasible	[66]
Radar altimeters	Infeasible	Dimensions
Scatterometers	Infeasible	Dimensions

Selva, D., Krejci, D., 2012. A survey and assessment of the capabilities of Cubesats for Earth observation. Acta Astronaut. 74, 50–68. <https://doi.org/10.1016/j.actaastro.2011.12.014>



Comparison

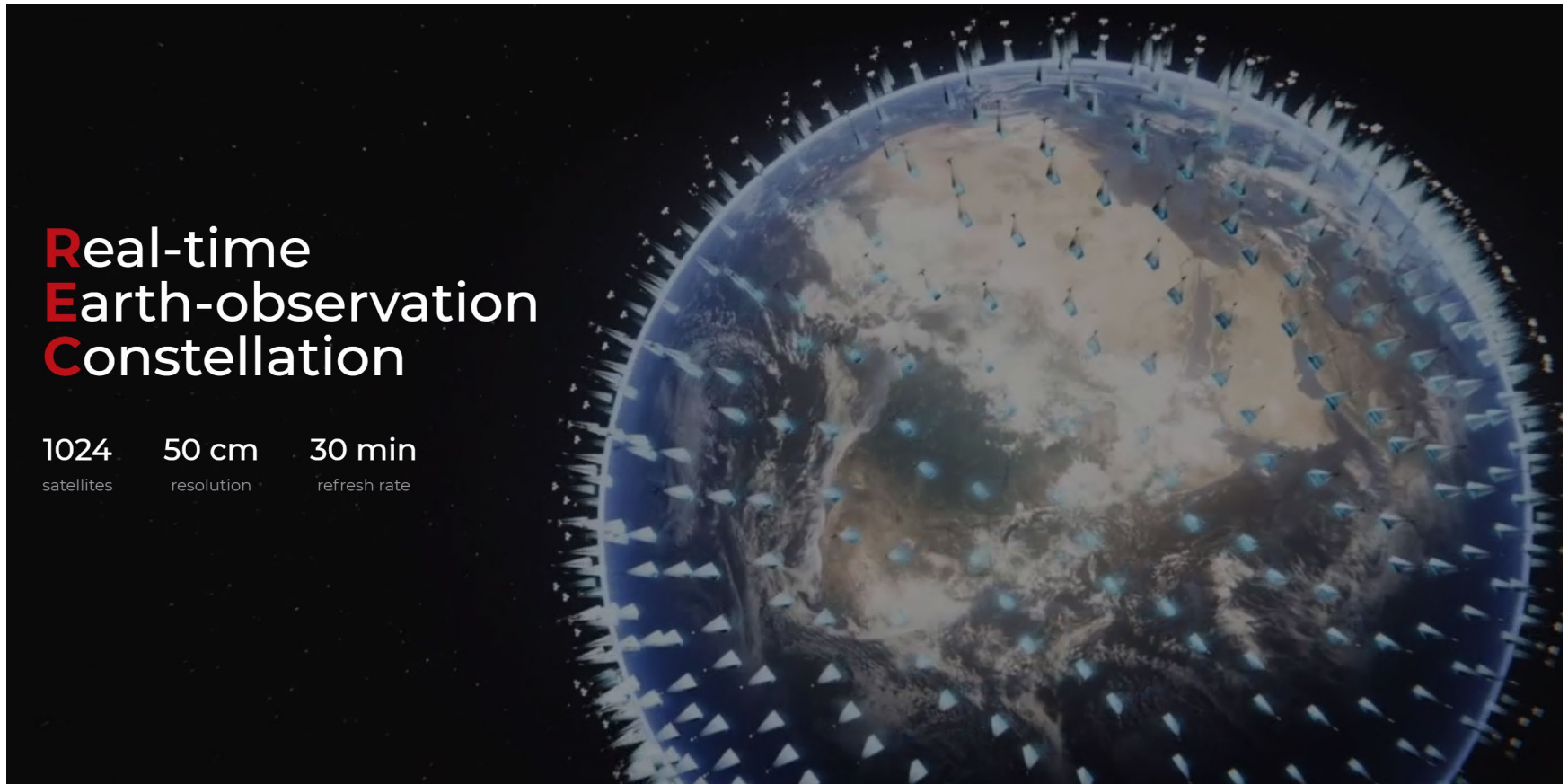


Planet, Planetscope
3.6m GSD, MS



WorldView 2,
approx. 2m, MS

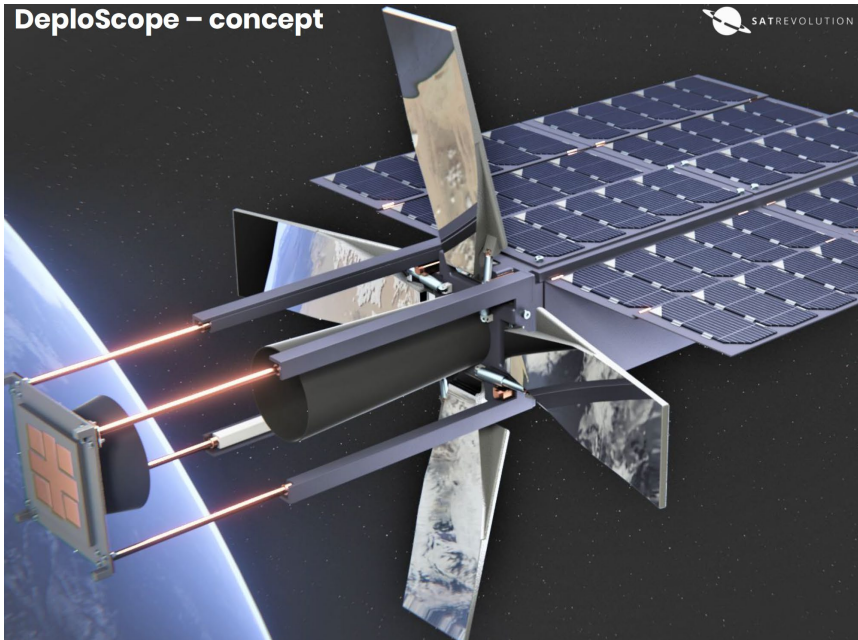
Satellite Revolution (?!)



Reference: <https://satrevolution.com/rec/>

Poźniak, Tomasz, 2019. Earth observation by utilizing nanosatellite constellation. Presented at the 11th European CubeSat Symposium, Luxembourg. <https://www.cubesatsymposium.eu/>

Satellite Revolution (?!)



No	Parameter	Value
1	Spectral range	VNIR – achromat
2	GSD (Ground Sample Distance)	1,1 m @350 km
3	Sensor size	8,5x7,1 mm
4	Pixel size	3,45x3,45 μ m
5	Focal length	1100 mm
6	f#	f/6,1
7	FoV (Field of View)	$\pm 0,288^\circ$
8	MTF (Modulation Transfer Function)	7% – 15,6% @93 lp/mm
9	Strehl ratio	0,8
10	Mass	>360 g (optics only)
11	Dimensions – stowed position	100x100x175 mm
12	Dimensions – deployed position	180x180x175 mm
13	Optical surface tolerances	P-V <1 μ m; RMS: 1 fringe @632,8 nm; focal length tolerance: 0,02%
14	Acceptable deployment error	7 μ m
15	Thickness of optical elements	3,5–6 mm
16	Materials TBD	Aluminium, SiC, Zerodur, Invar

Current stage

Phase 1 Q3 2019

- DeploScope's early design & NanoBus in orbit demonstration (as Światowid's base in mid July)

Phase 2 Q3 2021

- ScopeSat prototype in orbit demonstration

Phase 3 Q4 2022

- 16 ScopeSats in orbit – 16 h revisit time
- First constellation's iteration
- **Starting the satellite factory project**

Phase 4 Q4 2023

- 66 ScopeSats in orbit – 4 h revisit time
- Second constellation's iteration

Phase 5 Q2 2026

- 1024 ScopeSats in orbit – 30 min revisit
- Final constellation's iteration

Earth observation from LEO

‘The Risky Rush for Mega Constellations’



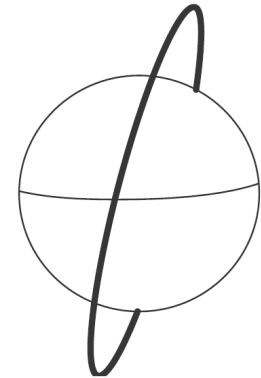
The European Space Agency recently was forced to maneuver its Aeolus satellite to avoid a potential collision with one of SpaceX's Starlink satellites. Credit: ESA and ATG medialab

Experts are alarmed by plans to launch tens of thousands of revolutionary satellites in coming years

O'Callaghan, Jonathan, 2019. The Risky Rush for Mega Constellations. Scientific American
<https://www.scientificamerican.com/article/the-risky-rush-for-mega-constellations/>

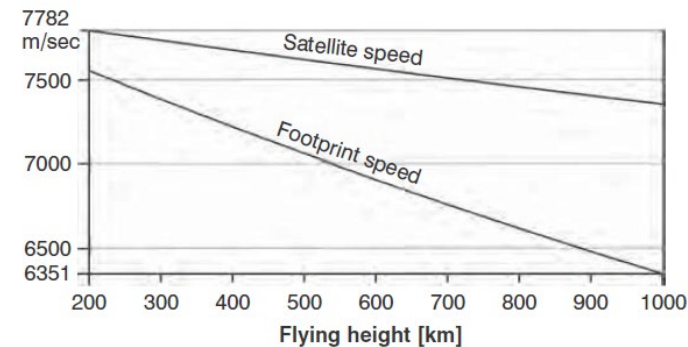
Earth observation from LEO

- Sun-synchronous, near-polar orbit with 98° inclination
 - most common orbit for optical EO satellites
 - orbit period approx. 90 min at 700-800km



- Satellite Speed as a function of flying height in a 'circular' orbit:

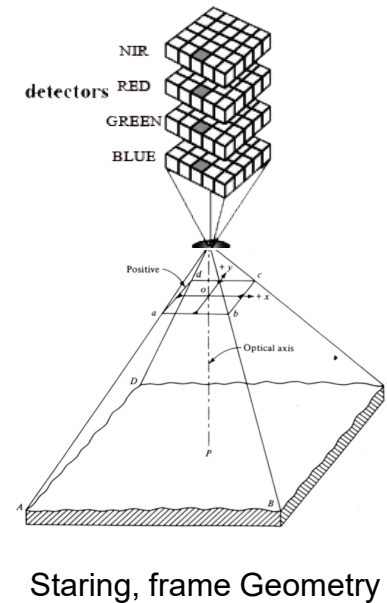
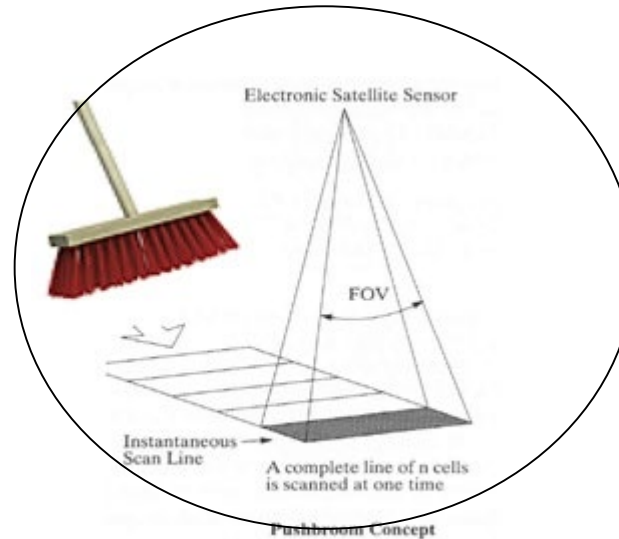
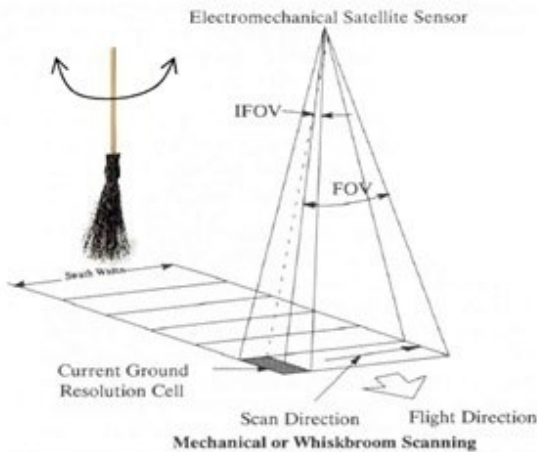
- Approx. 7 km/s for EO satellites
- t_{dwell} (1m GSD) $\sim 0.14\text{msec}$
- $t_{\text{int}} < t_{\text{dwell}}$



- Example Ikonos2 (typical for EO satellites)

- Inclination: 98.1°
- Period: 97 min
- Equatorial crossing: 10:30 am solar time
- Altitude: 681km

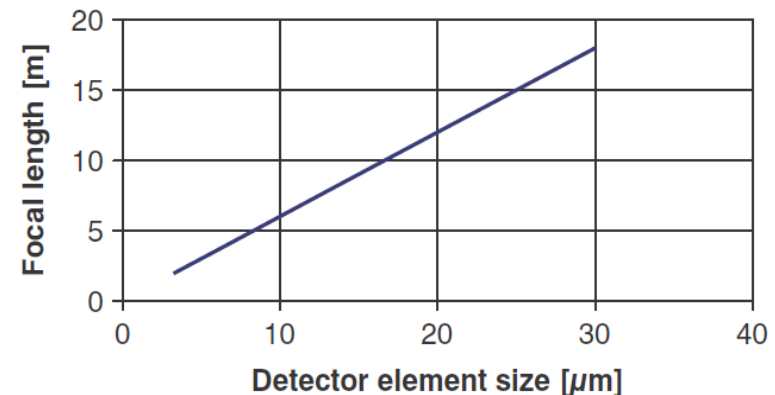
Satellite speed: 7.613 km/s
Footprint speed: 6.878 km/s



Relationship between detector element, focal length, orbit height and GSD:

$$f = \frac{H_{Orbit}}{GSD} \cdot x$$

The figure shows the relationship between required focal length and detector size for a orbit altitude of 600km at 1m GSD.



Requirements for Spaceborne Imaging Systems:

- Spatial resolutions and quality:

$$MTF_{SR} = MTF_{Optics} \cdot MTF_D \cdot MTF_{PS}$$

- Radiometric aspects:

- Higher resolution means smaller amounts of energy from smaller ground pixels
- **Time related** factor: dwell time (t_{dwell}) and **geometry related** factor (IFOV)
- E.g. the reduction of 10m to 1m GSD reduce the amount of energy at the detector by approx. 1000.

- Pointing accuracy:

- Start and stop pointing: < 500m
- Geolocation accuracy: 6.5m

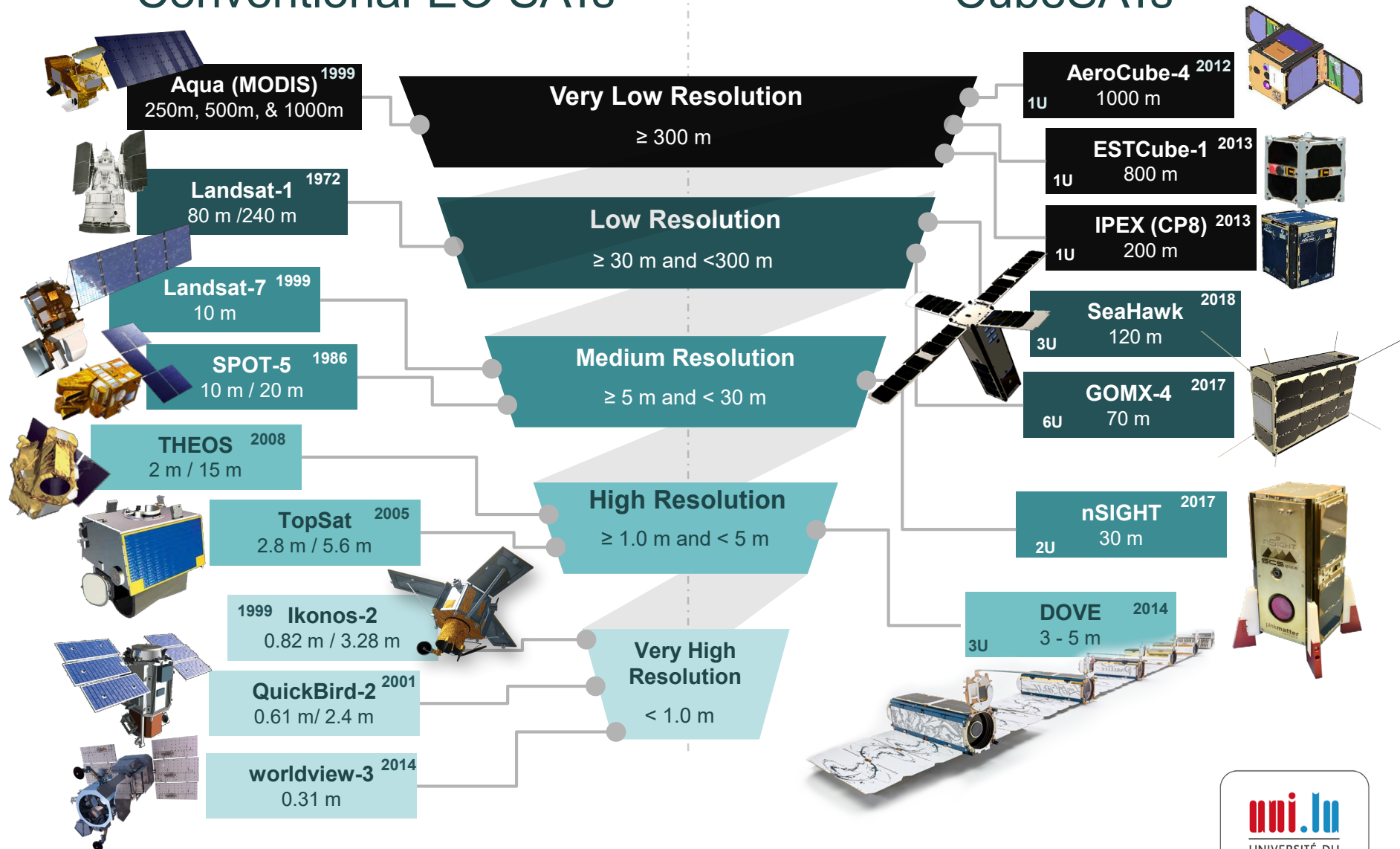
Common specification for high-resolutions EO systems

- Platform stability:

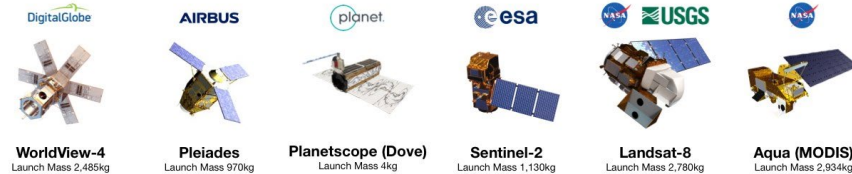
$$MTF_{PS} = MTF_{LM} \cdot MTF_J \cdot MTF_{sin}$$

Conventional EO SATs

CubeSATs

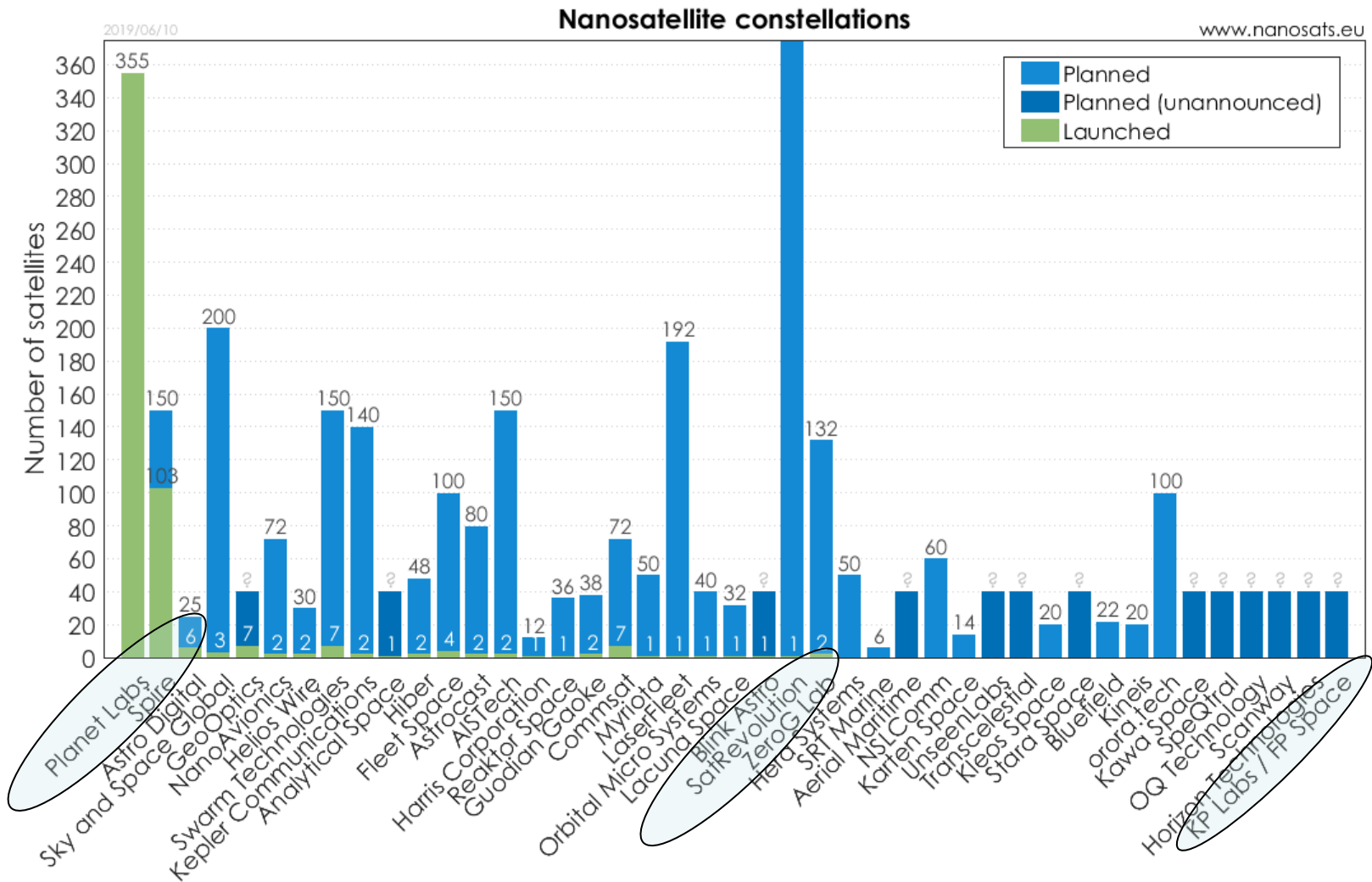


Evolution of Optical EO Satellites:



Launch Date	Organisation	Mission		Orbit	GSD	Sensor(s)	Pointing capability/Agility
1972 - 2013	NASA	Landsat	Landsat 1-3	907 to 915 km, 99°	80m	Multi spectral Scanner (MSS)	Up to 10.3° off nadir
			Landsat 4-5		30m	Thematic Mapper (TM)	
			Landsat 7	705 km, 98.2°	30m	Enhanced Thematic Mapper Plus (ETM+) 8-band whiskbroom scanning radiometer	Up to 7.5° off nadir
			Landsat 8		30m	Operational Land Imager (OLI) similar spectral bands to the ETM+	Up to 7.5° off nadir
1998-03-24	CNES (Centre national d'études spatiales)	Spot 1- 4		832 km, 98.8°	10 PAN / 20 MS	High-Resolution Visible and Infrared sensor (HRV IR)	± 27°
1999-09-24	Space Imaging/ GeoEye Inc.	Ikonos-2		681 to 709 km, 98.1°	1 m PAN (0.82 m at nadir), 4 m MS (3.2 m at nadir)	Kodak Optical Sensor Assembly (OSA) Pushbroom detector	±30°
1999-12-18	NASA	Terra	ASTER	705 km	15 m VNIR 30 m SWIR 90 m TIR 15 m Stereo	Advanced Spaceborne Thermal Emission and Reflection Radiometer (ASTER) 14 bands	0° / 27°
2002-05-04			MODIS	705 km	250 m (bands 1–2) 500 m (bands 3–7) 1000 m (bands 8–36)	Moderate Resolution Imaging Spectroradiometer (MODIS) Medium-resolution, multi-spectral, cross-track scanning radiometer 36 spectral bands	
2001-10-18	DigitalGlobe Inc	QuickBird-2		450 km, 97.2°	0.61 m (PAN) and at 2.4 m (MS)	Ball Global Imaging System 2000 (BGIS 2000) Pushbroom array	±30°
2002-05-04	CNES (Centre national d'études spatiales)	Spot-5		832 km, 98.7°	5m (single) 3.5m (double) PAN / 10m MS	High Resolution Geometric (HRG) High Resolution Stereo (HRS)	± 27° HRG ± 20° HRS
2008-08-29	RapidEye/Planet	RapidEye		630 km, 98°	6.5 m	Jena-Optronik RapidEye Earth Imaging System (REIS) Multispectral pushbroom sensor 5 spectral bands	±20°
2013-11-21	Skybox/Terra Bella/Planet	SkySat		600 km, 97.8°	90 cm PAN / 2.0 m MS	CMOS frame detectors (30f/s video from space)	
2014-08-13	DigitalGlobe Inc/MAXAR	WorldView-3		617 km	0.31 m PAN 1.24 m MS 3.7 m SWIR	Panchromatic, 8 Multispectral and 8 SWIR bands	±40° (nominal in any direction)
2015-06-23	ESA and EU (European Commission - Copernicus)	Sentinel-2 (a, b)		786 km, 98.5°	10 m: (VNIR) B2, B3, B4, B8 (4 bands) 20 m: B5, B6, B7, B8a, B11, B12 (6 bands) 60 m: B1, B9, B10 (3 bands)	Multispectral Imager (MSI) 13 bands VNIR + SWIR	

Some Planned and Launched CubeSat Missions

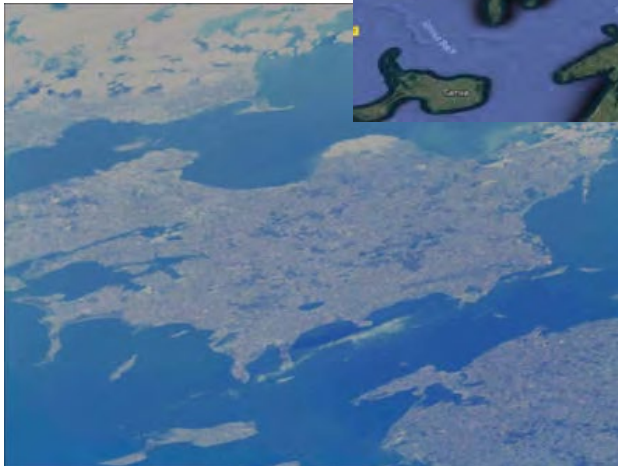
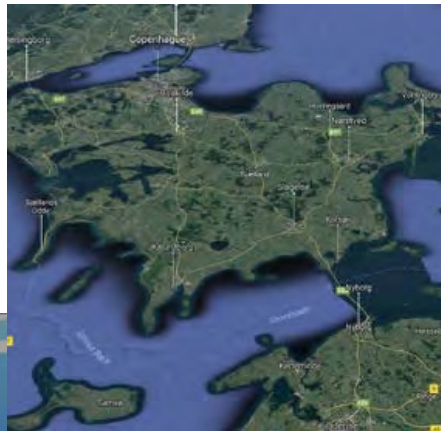


CubeSat Missions

	Launch Date	Organisation	Mission	Orbit	GSD	Sensor(s)	Resolution	Agility and Positioning
3U	2014 - 20XX	Planet Labs	Dove (Flock-xx xx)	400, 500, 600 (most are 500km)	3 - 5m	?		
	2015-12-16	Microspace Rapid Pte Ltd.	Athenoxat-1	540 km, 15 deg	Global view resolution: 1km Wide Angle resolution: 50m to 300m Narrow Beam Resolution: 1m to 20m	?	f/2 to f/10 Optics speed 2.5 deg narrow beam aperture Visible/IR, multiband, hyperspectral Spectrum up to 30Hz Video Refresh	ACS air-coil magnetorquers primarily for stabilization ADS sensors: coarse sun sensors, magnetometer & gyroscopes CDH & ADCS software including Nadir vector determination & payloads drivers
	2016-09-26	UK Space Agency	ALSAT-Nano	680 km, 98.2 deg, SSO		XCAM C3D2 (CMOS)	1200 x 1080 pixels Focal length: 45 cm	
	2018-12-03	University of North Carolina Wilmington	SeaHawk	580 km, 97.8 deg	120 m	push-broom design, with 4 linear array CCDs, each containing 3 rows of detectors	1800x 6000 pixels 8 bands deep	
6U	2017-08-14	NASA Jet Propulsion Laboratory	ASTERIA	400 km, 51.6 deg, ISS	30 m	Fairchild CIS2521 (CMOS)	2592 pixels x 2192 pixels Focal length: 85 mm Aperture diameter: 60.7 mm (f/1.4) Pixel size 6.5 µm x 6.5 µm Plate scale: 15.8 arcseconds per pixel Field of view: 11.2° x 9.6°	
	2018-02-02	GomSpace	GOMX-4	500 km	70 m	HyperScout camera from Cosine for hyperspectral images	4096 x 1850 pixels Spectral range: 400 - 1000 nm Spectral resolution: 15 nm Dynamic range: 12 bit SNR: 50-100	
	2019-09-05	ESA	Phi-Sat-1	450-550 km	VNIR 75 / TIR 390 (590km)	HyperScout-2, two spectral channels, each with 2D sensors operating in pushbroom mode.	FoV: VNIR 31° x 16° / TIR 31° x 16° Swath: VNIR 310 x 150 / TIR 310 x 150 Spectral bands: VNIR 45 / TIR 4 Spectral range [µm]VNIR 0.4 – 1.0 TIR 8.0 - 14	
	2022-12-31	KP Labs (FP Space)	intuition-1	?	?	hyperspectral instrument	Spectral resolution in the range of visible and near-infrared light The band is divided into 150 channels	

Example: GomX4 - Technology Demonstrator

GOMX-4A camera (GomSpaceNanoCam):



Sjaelland(DK)

Credit: GomSpace and Cosine; MarcoEsposito

GOMX-4B Hyperspectral Imager Hyperscout I by Cosine :

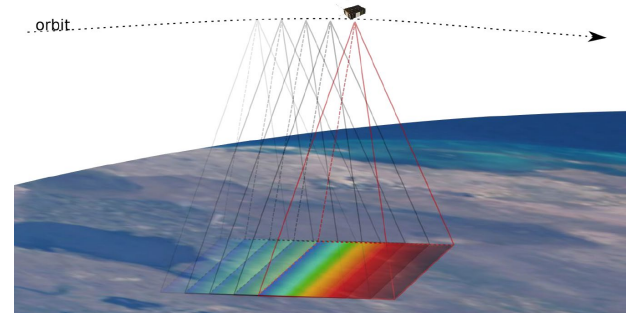


Figure 6: First light of HyperScout. False colour single image of the Scottish landscape between Glasgow and Edinburgh. Image acquired on the 20th of March 2018

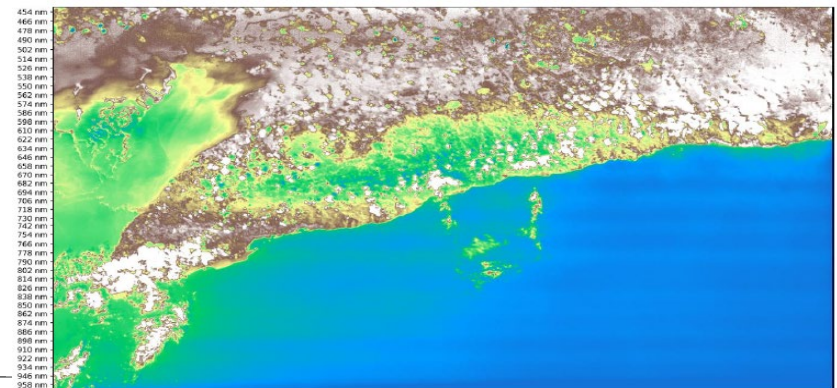


Figure 7: Commissioning image of HyperScout®. False colour single image of the southern Cuban coastline. Image acquired on the 26th of March 2018

Hyperscout applications

Disasters:

- Volcano and wild fire
- Flooding
- Monitoring geo hazards

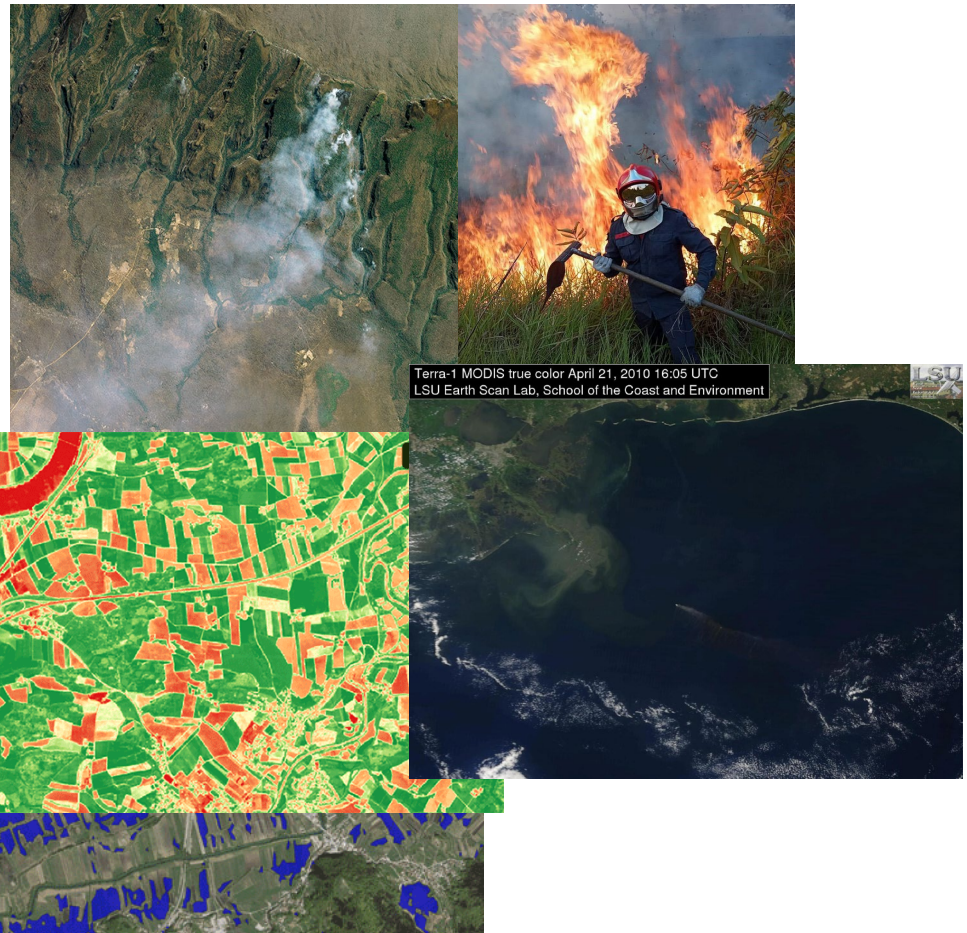
Environmental Monitoring:

- Water quality
- Pollution, oil spills etc.

Farming and agriculture:

- Crop monitoring
- Forest monitoring

Mapping



Sentinel-hub

Part 2: FDL-Europe Flood detection



DISASTER PREVENTION,
PROGRESS & RESPONSE
(FLOODS)

Flood Detection on Low-Cost Orbital Hardware

Team: Josh Veitch-Michaelis, Gonzalo Mateo-García, Lewis Smith, Silviu Oprea

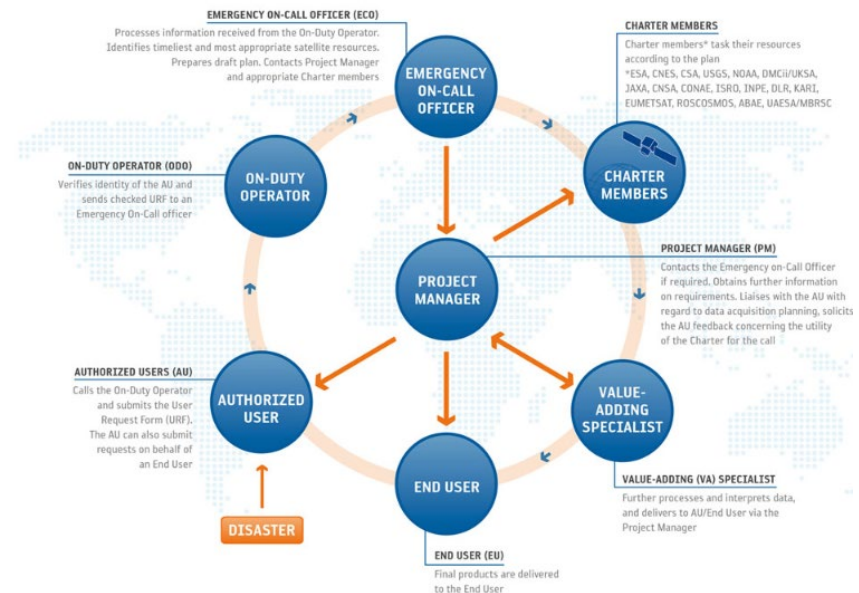
Mentors: Dietmar Backes, Atılım Güneş Baydin, Guy Schumann, Yarin Gal, Steve Reece

Flood detection and monitoring from Orbit



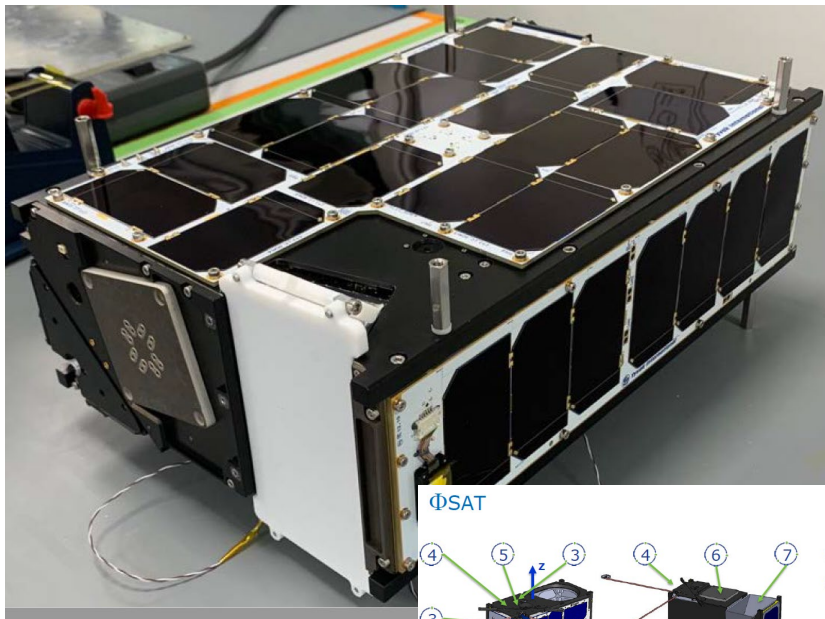
How can we reduce the time from disaster to data?

- Time for supply of data can be hours, on average 2 days: International “Space and Major Disasters” Charter
- How do we get data to the ground more quickly?
- How can we accelerate/automate the image analysis process?

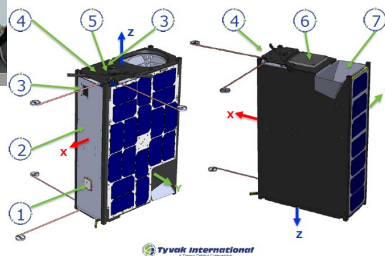


PhiSAT-1 On-board Processing

- FSSCat/PhiSat-1 technology demonstrator twin sat mission;
- Hyperspectral Sensor: Cosine Hyperscout II



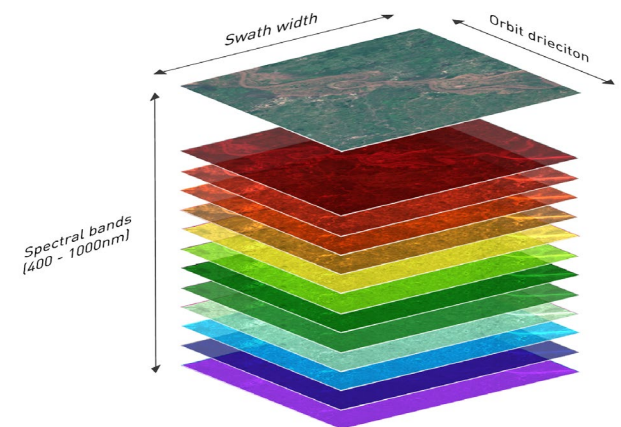
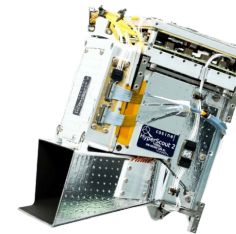
ΦSAT



Tyvak International

- 1 GPS Antenna
- 2 Vehicle BackPlane
- 3 XP/ZP Startracker Apertures
- 4 UHF Dipole Antenna Mechanism
- 5 Coarse Sun Sensor
- 6 S-Band Patch Antenna
- 7 HyperScout Telescope Baffle

0, 1, 1, 2, 3, 5, 8, 13, 21, 34, 7

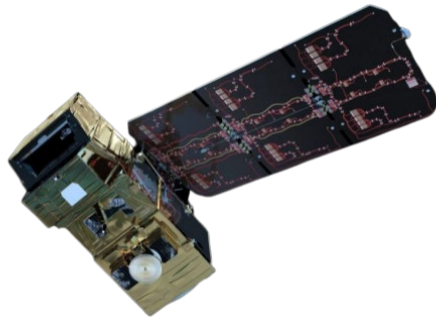


esa

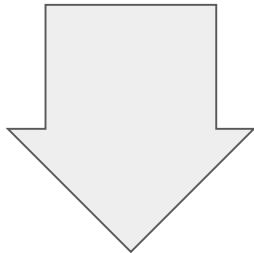
UNIVERSITÉ DU

Example: PhiSAT-1

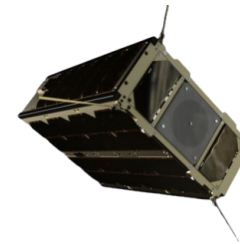
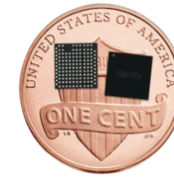
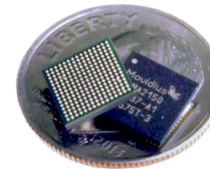
- Limited downlink capabilities
- AI processing on-board



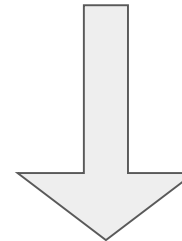
560 Mbps



ESA Maspalomas, Spain



< 1Mbps

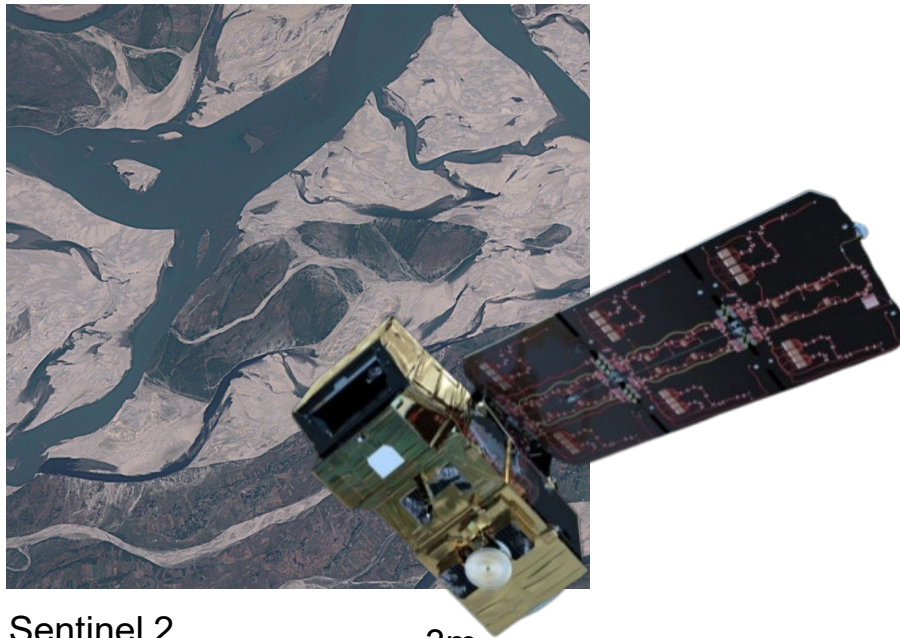


COTS ground station (ISIS)

Example: PhiSAT-1

Difference in Image Quality

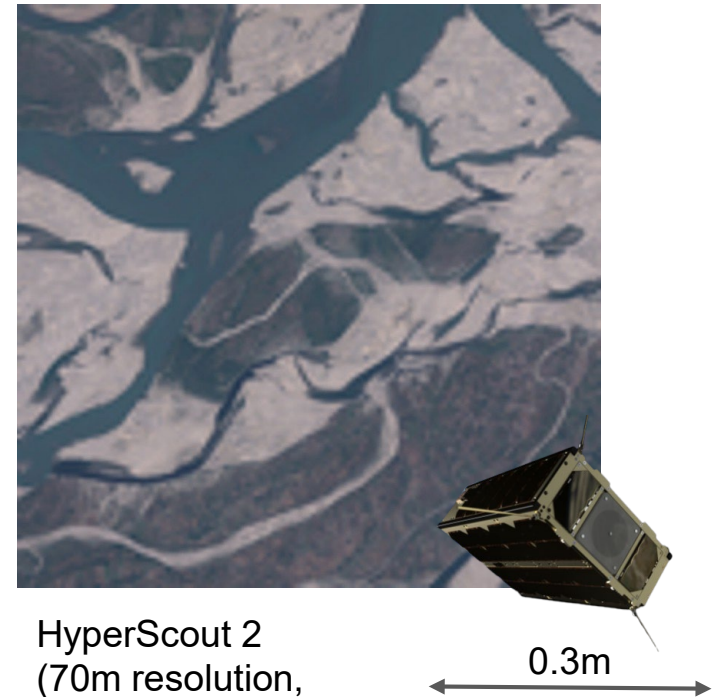
- Geometric, spectral, radiometric resolution,
- S/N ratio, motion blur etc.



Sentinel 2
(10m resolution)

3m

Antti Lipponen (Twitter @anttilip)



HyperScout 2
(70m resolution,
simulated)

0.3m

Approach

1. Dataset of global flood events and Sentinel 2 images
2. End to End machine learning models based on simulated imagery (degraded Sentinel 2 images)
3. Convert and test models on simulated images from Φ Sat-1 representative hardware (Intel Neural Compute Stick) prior to launch and deployment

Floods training data

- Combination of currently existing data bases
- Verification

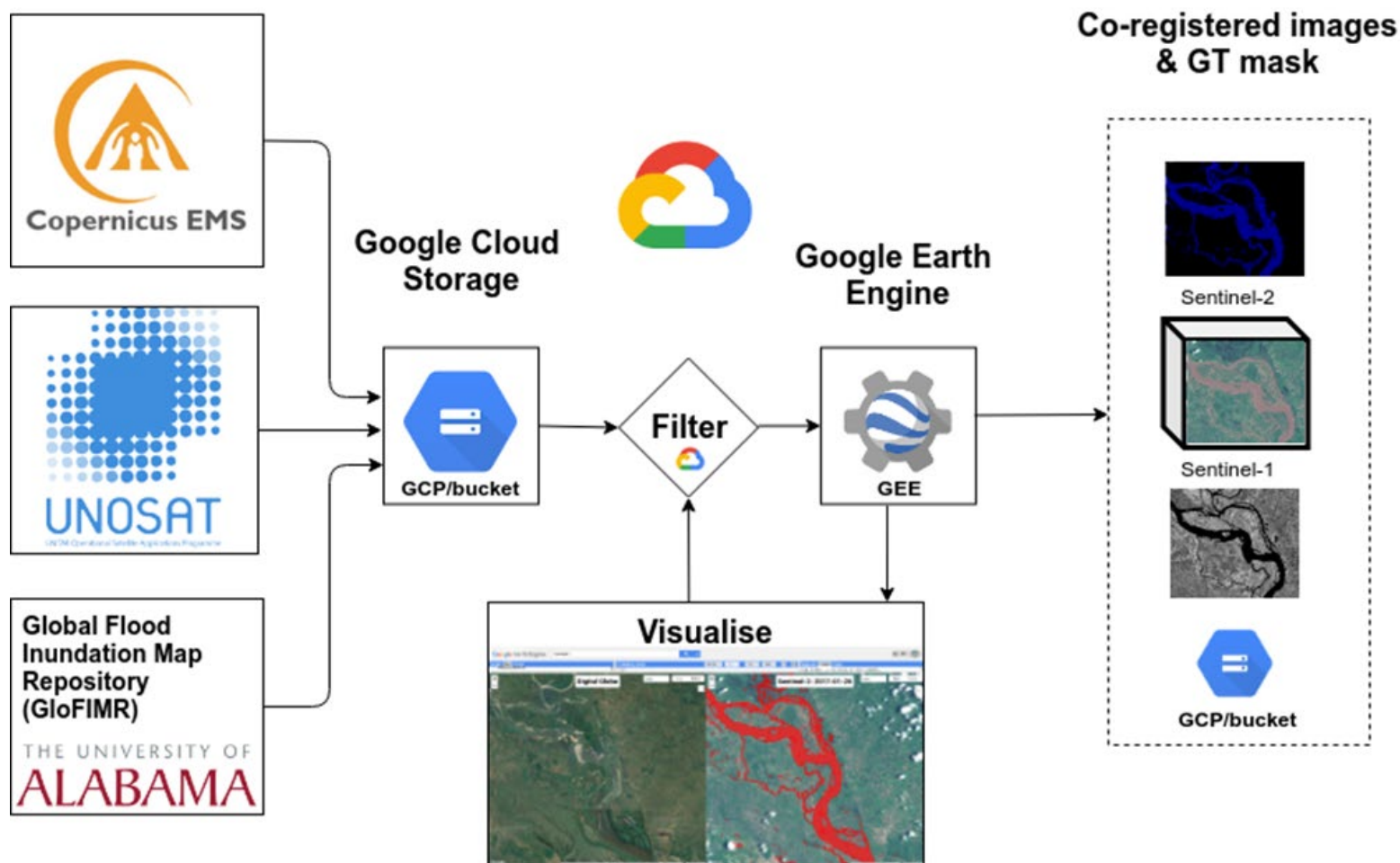
ML Flood detector

- Based on simulated training data
- Features:
 - Clouds
 - Flood water
- Model architecture

Optimisation for Satellite

- Sensor
- Onboard compute

WORLD FLOODS Dataset



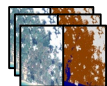
WORLDFLOODS Dataset



- 150 floods



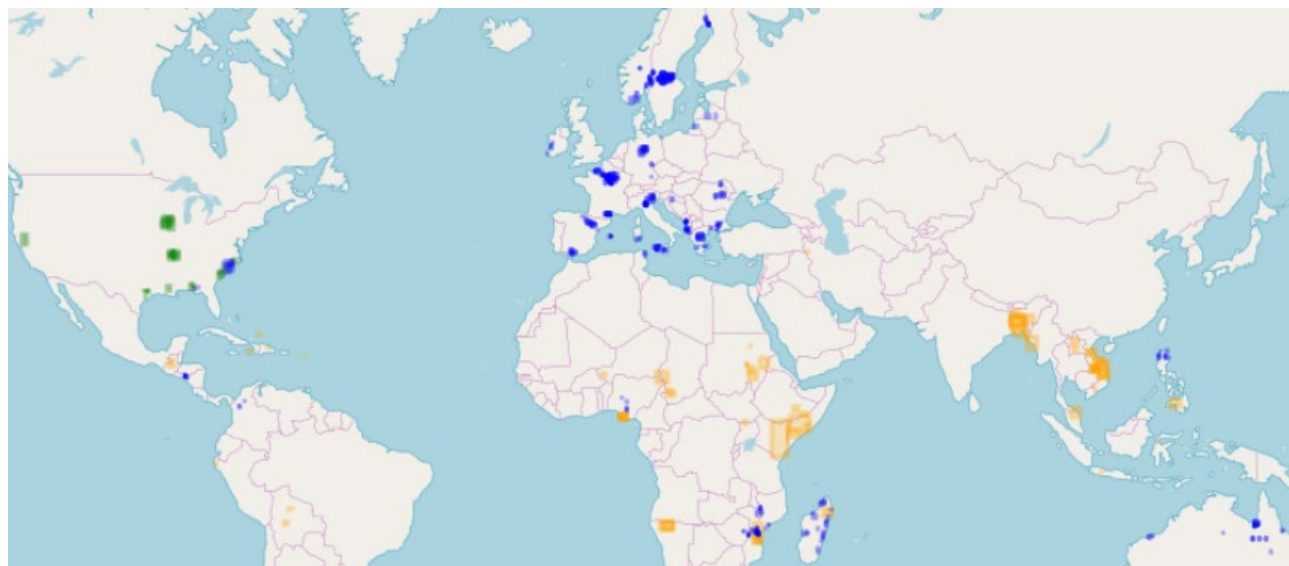
- 618 flood maps



235,000 patches
(256x256 px)



- 303 GB



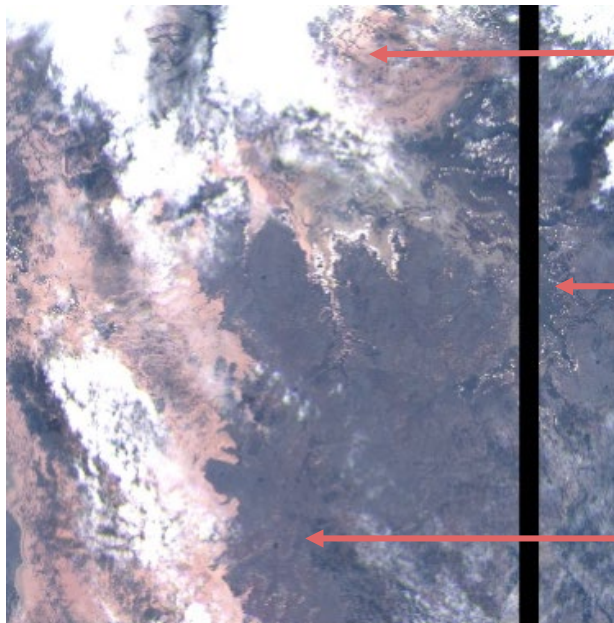
■ Unosat (127)

■ GloFIMR
(37)

■ Copernicus EMS (454)

WORLD FLOODS Dataset

Challenges

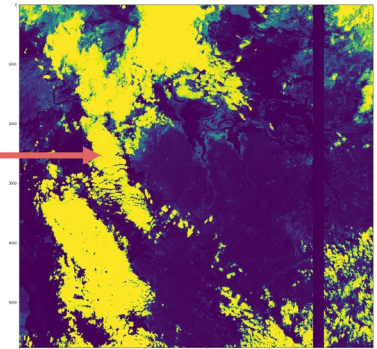


Clouds

Artifacts

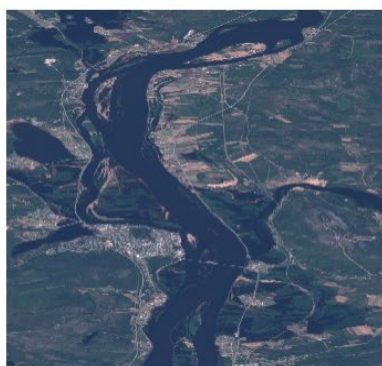
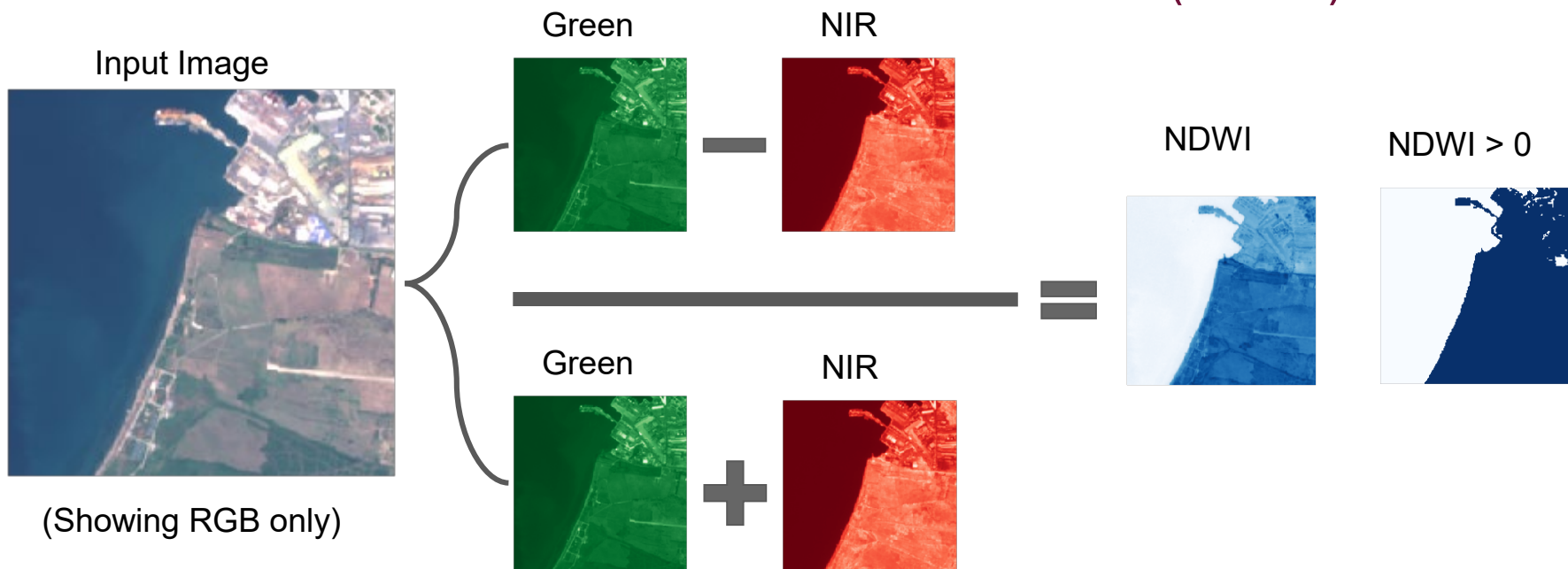
Image, label
noise

 cloudless github.com/sentinel-hub/sentinel2-cloud-detector



Type	%pixels
Clouds	45.9%
Land	43.7%
Invalid	7.5%
Water/Flood	2.85%

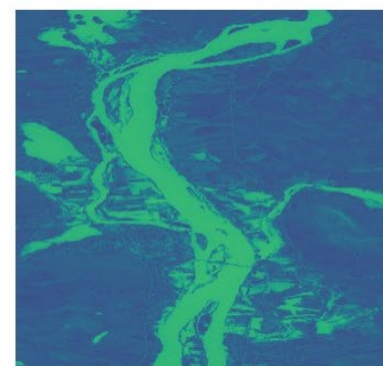
Baseline: Normalised Difference Water Index (NDWI)



Sentinel 2 RGB

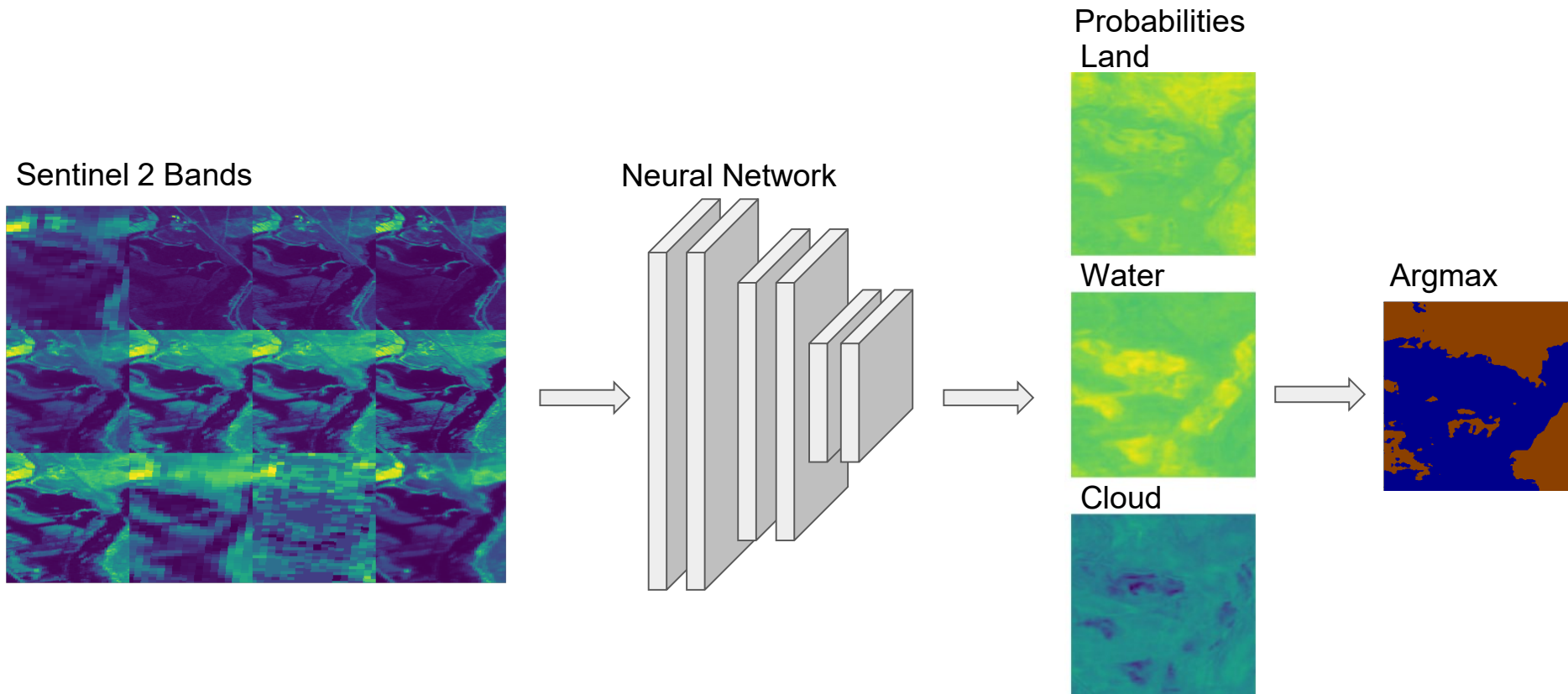


Ground truth



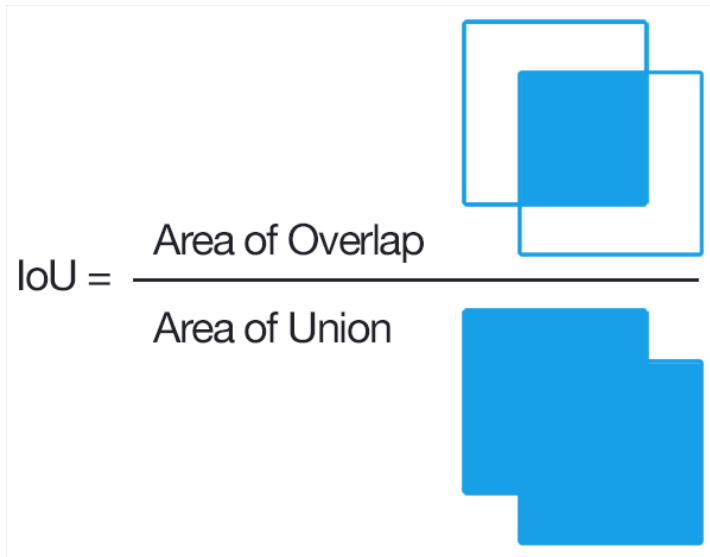
NDWI

Model architecture: fully convolutional neural networks

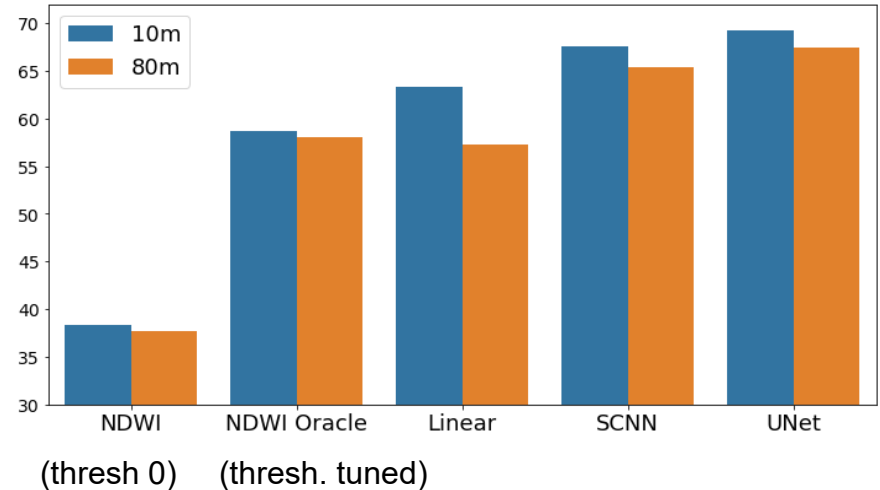


End to End machine learning approach

Evaluation



We experiment with training on both raw (10m/pixel) and degraded (80m/pixel) images



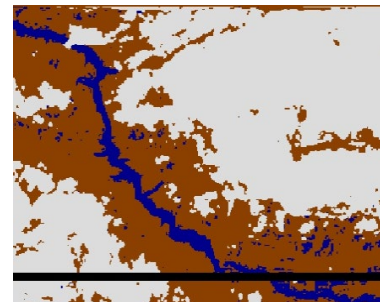
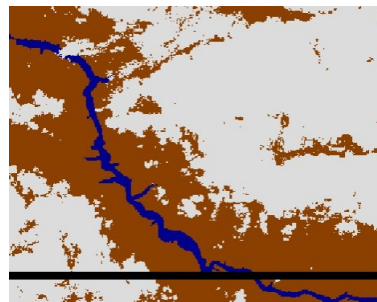
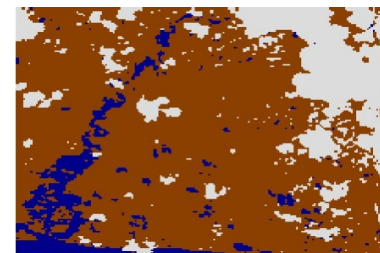
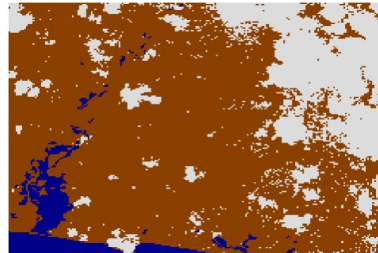
	IoU		Recall	
	10m	80m	10m	80m
NDWI (thresh 0)	38.37	37.47	43.7	43.13
NDWI (thresh. tuned)	58.63	58.09	93.26	97.72
Linear	63.34	57.25	96.24	95.24
SCNN	67.54	65.42	97.43	97.03
UNet	69.25	67.51	97.34	96.56



Simulation on Satellite hardware

Optimisation: 'lean' deep learning algorithm to be deployed directly on the satellite

- Performance drop < 1%
- Deep learning NN: Model size < 0.5 MB
- 12MP image mapped in < 1 minute



Outlook

- Image resolution quality of conventional EO Satellites will remain superior for the foreseeable future
- CubeSat technology will mature rapidly; technology demonstrators will soon become operational systems
- AI on small CubeSats satellites is expected to be a game changer (for some applications)
 - We hope that our Flood detector will be tested onboard Small or Nano satellites in the future

Thanks for listening!



RSS-Hydro, S.à r.l.-S



UNIVERSITÉ DU
LUXEMBOURG

DISASTER PREVENTION, PROGRESS & RESPONSE (FLOODS)

Flood Detection on Low-Cost Orbital Hardware

Team: Josh Veitch-Michaelis, Gonzalo Mateo-García, Lewis Smith, Silviu Oprea

Mentors: Dietmar Backes, Atılım Güneş Baydin, Guy Schumann, Yarin Gal, Steve Reece



Low Earth Orbits (LEO)

- Revisit time:
 - is a function of swath width, spacecraft agility/pointability and the number of space crafts
 - Often called ‘temporal resolution’

